

Topic 2 - Stronger Eyes and Better Numbers

Telescopes

<http://www.nasa.gov/NAS/SpaceSettlement/>

Telescopes allow us to see objects that are very distant in space.

Optical Telescopes <http://cdsweb.u-strasbg.fr/astroweb/optical.html>

In 1608, Hans Lippershey made one of the first telescopes – but it was Galileo Galilei who made practical use of it. The observations he made included:

- The moon had blemishes (mountains and craters like the Earth).
- Sun spots indicated that it rotates on its axis.
- Jupiter's moons orbit the planet.
- Planets were disk-shaped, but because the stars were still pinpoints, they were further away.

Galileo's Approach to Inquiry

Galileo's observations supported Copernicus's Sun-centered model but not Ptolemy's Earth-centered model. The reason for his beliefs was that, the moons he observed orbiting Jupiter, indicated that the earth was not the centre of the universe.

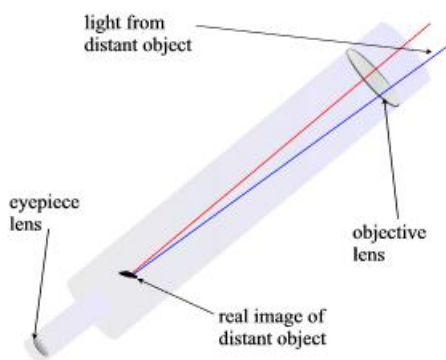
Resolving Power

Resolving power is the kind of power that produces finer detail of the object being viewed because of the diameter of the objective lens. There is a limit to the size of lens that a refracting telescope can have. Diameters over 1 meter will cause the lens to warp.

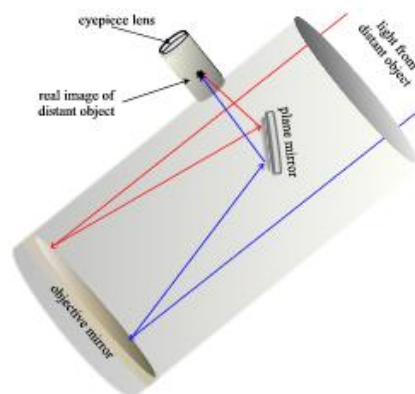
Refractors and Reflectors

Optical telescopes are 'light collectors'. The series of lenses or mirrors enable the optical device to collect and focus the light from stars. There are two types of optical telescopes:

The first telescope designed was a simple **refracting telescope**. It uses two **lenses** to gather and focus starlight



Reflecting telescopes use **mirrors** instead of lenses to gather and focus the light from the stars. A process called '**spin-casting**' today makes mirrors, by pouring molten glass into a spinning mould. The glass is forced to the edges, cooled and solidified. Mirrors as large as 6m across have been made using this method.



Review Optical Telescopes Notes **Grade 8 Science Focus****Topic 5: Extending Human Vision** <http://www.edquest.ca/content/view/188/>

An innovation for ground-based optical reflecting telescopes is the use of **segmented mirrors** (a segmented-mirror telescope uses several lightweight-segments to build one large mirror).

Interferometry: Combining Telescopes for Greater Power. The technique of using a number of telescopes in combination is called **interferometry**. When working together, these telescopes can detect objects in space with better clarity and at greater distances than any current Earth-based observatory.

http://www.space.com/scienceastronomy/astronomy/interferometry_101.html

To improve the views of space, astronomers are able to access images from a telescope in space. Free from the interferences of weather, clouds humidity and even high winds, the **Hubble Space Telescope**, launched in 1990, orbits 600 kms above the Earth, collecting images of extremely distant objects. It is a cylindrical reflecting telescope, 13 m long and 4.3 m in diameter. It is **modular** (parts can be removed and replaced) and is serviced by shuttle astronauts.

Hubble Facts: http://hubblesite.org/reference_desk/facts_and_figures/

Copernicus's Sun-centred Revolution Continues

Although Galileo's observations helped to confirm the Sun-centered model of the universe, it was Johannes Kepler who solved the remaining problem of the epicycles. To do this, his calculations insisted that the orbits of the planets should be **elliptical**, instead of circular. An ellipse is a figure that looks like a squashed circle.

Tycho Brahe was an observation genius in astronomy before the age of the telescope. The mural, or Tychonian, quadrant was actually a very large brass quadrant, affixed to a wall. Its radius measured almost two meters and was graduated in tens of seconds. Sightings were taken along the quadrant through the small window in the opposing wall, to which Tycho points. The clock shown at the bottom right, accurate to seconds, allowed the observers to note the precise moment of observation.



With access to Tycho Brahe's star charts, Kepler mathematically worked out the orbit of Mars and found that it only worked if the orbit was elliptical. He also figured out the shape and scale of the entire known solar system.

Universal Gravitation

Isaac Newton stated the law of universal gravitation eighty years after Kepler's contribution about elliptical orbits of the planets. Newton's law states that there is a gravitational force between all objects that pulls them together.

An orbit is the result of the attractive force of gravity balancing the straightforward movement of a planet because of velocity.

